

Comment on “Self-interacting Elko dark matter with an axis of locality”

Edmundo Capelas de Oliveira* and Waldyr Alves Rodrigues, Jr.†

 Institute of Mathematics, Statistics and Scientific Computation, IMECC-UNICAMP,
 13083-859 Campinas, São Paulo, Brazil

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In this comment we show that the statement in Ahluwalia *et al.* [Phys. Rev. D **83**, 065017 (2011)] that the existence of Elko spinor fields imply in an *axis of locality* is equivocated. The anticommutator $\{\Lambda(\mathbf{x}, t), \Pi(\mathbf{x}, t)\}$ is strictly local.

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**I. THE INTEGRAL APPEARING
 IN THE PROPAGATOR**

In Ref. [1] the authors calculated the propagator for an Elko spinor field supposed to satisfy the Klein-Gordon equation in Minkowski spacetime. They obtained a sum of two terms, the first being the usual propagator (fundamental solution) for a Klein-Gordon field and the second involves evaluation of the integral [see their Eq. (6.20)]

$$\int \frac{d^4 p}{(2\pi)^4} e^{-ip_\mu(x'^\mu - x^\mu)} \frac{i\bar{\omega}}{p_\mu p^\mu - m^2 + i\epsilon} \mathcal{G}(\mathbf{p}). \quad (1)$$

The calculation is done in an inertial reference frame¹ $\mathbf{e}_0 = \partial/\partial t$ with arbitrary spatial axes $\langle \mathbf{e}_1 = \frac{\partial}{\partial x}, \mathbf{e}_2 = \frac{\partial}{\partial y}, \mathbf{e}_3 = \frac{\partial}{\partial z} \rangle$ chosen in such a way that together with \mathbf{e}_0 defines a global orthonormal tetrad in Minkowski spacetime. We next introduce spherical coordinates associated with the selected orthonormal triad $\langle \mathbf{e}_1, \mathbf{e}_2, \mathbf{e}_3 \rangle$ and write

$$\mathbf{p} = (r \sin\theta \cos\varphi, r \sin\theta \sin\varphi, r \cos\theta).$$

 Then²

$$\mathcal{G}(\mathbf{p}) := \gamma^5 \gamma^\mu n_\mu, \quad (2)$$

where the spacelike vector field $n = n^\mu \mathbf{e}_\mu$ is

$$n_\mu := (0, \mathbf{n}),$$

$$\mathbf{n} := \frac{1}{\sin\theta} \frac{\partial}{\partial \varphi} \left(\frac{\mathbf{p}}{|\mathbf{p}|} \right) = (-\sin\varphi, \cos\varphi, 0).$$

Then the authors claim:

If there is no preferred direction, and since we are integrating over all momenta, we are free to choose a coordinate system in which $\mathbf{x}' - \mathbf{x}$ lies in the \hat{z} direction. In this special case, the $\mathbf{p} \cdot (\mathbf{x}' - \mathbf{x})$ depends only on

$p(=|\mathbf{p}|)$ and θ , but not on φ . Thus, the only φ -dependence in the whole integrand comes from \mathcal{G} which depends on φ in such a manner that an integral over one period vanishes.

Remark 1.—It is very important to remark that if this integral would result as non-null, the fundamental solution for the Klein-Gordon operator would have an additional term, something that obviously cannot be the case.

II. THE INTEGRAL OF $\mathcal{G}(\mathbf{p})$

On the other hand in Ref. [4] authors calculated the anticommutator of an Elko spinor field with its canonical momentum getting their Eq. (42), i.e.,

$$\{\Lambda(\mathbf{x}, t), \Pi(\mathbf{x}, t)\} = i\delta(\mathbf{x} - \mathbf{x}') + i \int \frac{d^3 p}{(2\pi)^3} e^{i\mathbf{p} \cdot (\mathbf{x} - \mathbf{x}')} \mathcal{G}(\mathbf{p}). \quad (3)$$

There they claim:

“Since the integral on the right hand side of Eq. (42) vanishes only along the $\pm \hat{z}_e$ axis, the preferred axis also becomes the axis of locality.”

Let us examine if that claim is correct. Call $|\mathbf{x} - \mathbf{x}'| = \Delta$, and put

$$(\mathbf{x} - \mathbf{x}') = \Delta(\sin\theta_\Delta \cos\varphi_\Delta, \sin\theta_\Delta \sin\varphi_\Delta, \cos\theta_\Delta). \quad (4)$$

Calculation of the integral in the second member of Eq. (3) resumes in the calculation of the following integrals³

$$\mathbf{J}(\Delta) = \int_0^\infty dr r^2 \int_0^\pi d\theta \sin\theta \int_0^{2\pi} d\varphi e^{ir\Delta f(\theta, \theta_\Delta, \varphi, \varphi_\Delta)} \sin\varphi, \quad (5)$$

and

$$\mathbf{K}(\Delta) = \int_0^\infty dr r^2 \int_0^\pi d\theta \sin\theta \int_0^{2\pi} d\varphi e^{ir\Delta f(\theta, \theta_\Delta, \varphi, \varphi_\Delta)} \cos\varphi, \quad (6)$$

³We take $\Delta \neq 0$. For $\Delta = 0$ it is obvious that $\mathbf{J}(0) = \mathbf{K}(0) = 0$.

*capelas@ime.unicamp.br

†walrod@ime.unicamp.br

¹In relativity theory reference frames are represented by time-like vector fields \mathbf{Z} on the manifold modeling spacetime. In special relativity if D is the Levi-Civita connection of the Minkowski metric η , an inertial frame is a timelike vector field \mathbf{I} such that $D\mathbf{I} = 0$. Details can be found, e.g., in Refs. [2,3].

²See also Eq. (31) in Ref. [4].

with $f(\theta, \theta_\Delta, \varphi, \varphi_\Delta) = \sin\theta \cos\varphi \sin\theta_\Delta \cos\varphi_\Delta + \sin\theta \sin\varphi \times \sin\theta_\Delta \sin\varphi_\Delta + \cos\theta \cos\theta_\Delta$.

We will now calculate the integrals in Eqs. (5) and (6) in the cases when $\mathbf{x} - \mathbf{x}'$ lies, respectively, in the \mathbf{e}_1 , \mathbf{e}_2 , and \mathbf{e}_3 directions. We will call the respective integrals $\mathbf{J}_i(\Delta)$ and $\mathbf{K}_i(\Delta)$, for $i = 1, 2, 3$.

We start with the observation that it is trivial to verify that $\mathbf{J}_3(\Delta) = 0$ and $\mathbf{K}_3(\Delta) = 0$. To continue we calculate $\mathbf{J}_2(\Delta)$.

So, let us choose the spatial axis such that $(\mathbf{x} - \mathbf{x}') = \Delta \mathbf{e}_2 = \Delta(0, 1, 0)$ and perform the nontrivial exercise of calculating the value of the integral given by Eq. (5) in this case, i.e.,

$$\mathbf{J}_2(\Delta) = \int_0^\infty dr r^2 \int_0^\pi d\theta \sin\theta \int_0^{2\pi} d\varphi \sin\varphi e^{ir\Delta \sin\theta \sin\varphi}. \quad (7)$$

We start evaluating the φ -integral,

$$\Omega(\xi) = \int_0^{2\pi} d\varphi \sin\varphi e^{i\xi \sin\varphi}, \quad (8)$$

where $\xi := r\Delta \sin\theta$. Observe that $\Omega(\xi) = -id\Lambda(\xi)/d\xi$ where (see 8.473-4, page 968 of Ref. [5])

$$\begin{aligned} \Lambda(\xi) &= \int_0^{2\pi} d\varphi e^{i\xi \sin\varphi} \\ &= 2 \int_0^\pi d\varphi \cos(r\Delta \sin\theta \sin\varphi) \\ &= 2\pi J_0(r\Delta \sin\theta), \end{aligned} \quad (9)$$

with J_0 the zero order Bessel function. So,

$$\Omega(\xi) = 2\pi i J_1(r\Delta \sin\theta), \quad (10)$$

where J_1 is the first-order Bessel function. Next we evaluate

$$\Xi(r, \Delta) = 2\pi i \int_0^\pi d\theta \sin\theta J_1(r\Delta \sin\theta). \quad (11)$$

Using the relation 6.681-8, page 739 of Ref. [5], namely

$$\int_0^\pi dx \sin(2\mu x) J_{2\nu}(2a \sin x) = \pi \sin(\mu\pi) J_{\nu-\mu}(a) J_{\nu+\mu}(a), \quad (12)$$

valid for $\text{Re}(\nu) > -1$ we see that identifying $2\mu = 1$, $2\nu = 1$, and $2a = r\Delta$ we can write

$$\Xi(r, \Delta) = 2\pi^2 i J_0\left(\frac{r\Delta}{2}\right) J_1\left(\frac{r\Delta}{2}\right). \quad (13)$$

So, putting $t = r\Delta/2$ we have

$$\mathbf{J}_2(\Delta) = \frac{16\pi^2 i}{\Delta^3} \int_0^\infty dt t^2 J_0(t) J_1(t). \quad (14)$$

Now, recall relation 6.626-3, page 715 of Ref. [5] (with $\beta = 1$), namely

$$\int_0^\infty dx e^{-2\alpha x} x J_0(x) J_1(x) = \frac{1}{2\pi} \left[\frac{K\left(\frac{1}{\sqrt{1+\alpha^2}}\right) - E\left(\frac{1}{\sqrt{1+\alpha^2}}\right)}{\sqrt{1+\alpha^2}} \right]. \quad (15)$$

Then we see that

$$\mathbf{J}_2(\Delta) = -\frac{4\pi i}{\Delta^3} \lim_{\alpha \rightarrow 0} \frac{d}{d\alpha} \left[\frac{K\left(\frac{1}{\sqrt{1+\alpha^2}}\right) - E\left(\frac{1}{\sqrt{1+\alpha^2}}\right)}{\sqrt{1+\alpha^2}} \right]. \quad (16)$$

Recalling relations 8.113 and 8.114, page 905 of Ref. [5], we have for the elliptic functions K and E

$$\begin{aligned} K\left(\frac{1}{\sqrt{1+\alpha^2}}\right) &= \frac{\pi}{2} {}_2F_1\left(\frac{1}{2}, \frac{1}{2}; 1; \frac{1}{1+\alpha^2}\right), \\ E\left(\frac{1}{\sqrt{1+\alpha^2}}\right) &= \frac{\pi}{2} {}_2F_2\left(-\frac{1}{2}, \frac{1}{2}; 1; \frac{1}{1+\alpha^2}\right), \end{aligned} \quad (17)$$

where ${}_2F_1$ are Gauss hypergeometric functions. So,

$$\begin{aligned} \int_0^\infty dx e^{-2\alpha x} x J_0(x) J_1(x) &= \frac{1}{4} \left[\frac{{}_2F_1\left(\frac{1}{2}, \frac{1}{2}; 1; \frac{1}{1+\alpha^2}\right) - {}_2F_1\left(-\frac{1}{2}, \frac{1}{2}; 1; \frac{1}{1+\alpha^2}\right)}{\sqrt{1+\alpha^2}} \right] \end{aligned} \quad (18)$$

and to end our long calculation we must evaluate the limit when $a \rightarrow 0$ of

$$\frac{d}{d\alpha} \left[\frac{{}_2F_1\left(\frac{1}{2}, \frac{1}{2}; 1; \frac{1}{1+\alpha^2}\right) - {}_2F_1\left(-\frac{1}{2}, \frac{1}{2}; 1; \frac{1}{1+\alpha^2}\right)}{\sqrt{1+\alpha^2}} \right]. \quad (19)$$

Recalling that (see, e.g., page 281 of Ref. [6])

$$\frac{d}{dz} {}_2F_1(a, b; c; z) = \frac{ab}{c} {}_2F_1(a+1, b+1; c+1; z),$$

and that ${}_2F_1(a, b; c; z) = \sum_{n=0}^\infty \frac{(a)_n (b)_n}{(c)_n} \frac{z^n}{n!}$ (with $(a)_n$, $(b)_n$, $(c)_n$ the Pochhammer symbols) we see that we are justified in writing

$$\begin{aligned} \lim_{\alpha \rightarrow 0} \frac{d}{d\alpha} \left[\frac{{}_2F_1\left(\frac{1}{2}, \frac{1}{2}; 1; \frac{1}{1+\alpha^2}\right)}{\sqrt{1+\alpha^2}} \right] &= \lim_{\alpha \rightarrow 0} \left\{ \begin{aligned} &-\alpha(1+\alpha^2)^{-3/2} {}_2F_1\left(\frac{1}{2}, \frac{1}{2}; 1; \frac{1}{1+\alpha^2}\right) \\ &-2\alpha(1+\alpha^2)^{-5/2} {}_2F_1'\left(\frac{1}{2}, \frac{1}{2}; 1; \frac{1}{1+\alpha^2}\right) \end{aligned} \right\} = 0. \end{aligned} \quad (20)$$

Also,

$$\lim_{\alpha \rightarrow 0} \frac{d}{d\alpha} \left[(1+\alpha^2)^{-1/2} {}_2F_1\left(-\frac{1}{2}, \frac{1}{2}; 1; \frac{1}{1+\alpha^2}\right) \right] = 0. \quad (21)$$

Finally, using Eqs. (20) and (21) in Eq. (16) we have that $\mathbf{J}_2(\Delta) = 0$.

We now evaluate

$$\mathbf{K}_2(\Delta) = \int_0^\infty dr r^2 \int_0^\pi d\theta \sin\theta \int_0^{2\pi} d\varphi e^{ir\Delta \sin\theta \sin\varphi} \cos\varphi. \quad (22)$$

Calling $\mathbf{x} = \sin\varphi$ and $r\Delta \sin\theta = \alpha$ we see that (taking into account relation 3.715-9, page 401 of Ref. [5] with $n = 1$) we have

$$\begin{aligned} \int_0^{2\pi} d\varphi e^{ir\Delta \sin\theta \sin\varphi} \cos\varphi &= \int_{-\pi}^{\pi} d\varphi e^{ir\Delta \sin\theta \sin\varphi} \cos\varphi \\ &= -2 \int_0^\pi d\mathbf{x} \cos\varphi \cos(\mathbf{x} \sin\varphi) \\ &= -2[1 + (-1)] \frac{\pi}{2} J_1(\mathbf{x}) = 0 \end{aligned} \quad (23)$$

and we conclude that $\mathbf{K}_2(\Delta) = 0$.

We now evaluate $\mathbf{J}_1(\Delta)$ and $\mathbf{K}_1(\Delta)$. Observe that

$$\int_0^{2\pi} d\varphi \sin\varphi e^{ir\Delta \sin\theta \cos\varphi} = - \int_{-\pi}^{\pi} d\varphi \sin\varphi e^{ir\Delta \sin\theta \cos\varphi} = 0, \quad (24)$$

since the integrand is an odd function and the interval is symmetric. So, we have

$$\mathbf{J}_1(\Delta) = \int_0^\infty dr r^2 \int_0^\pi d\theta \sin\theta \int_0^{2\pi} d\varphi \sin\varphi e^{ir\Delta \sin\theta \cos\varphi} = 0. \quad (25)$$

It remains to evaluate

$$\mathbf{K}_1(\Delta) = \int_0^\infty dr r^2 \int_0^\pi d\theta \sin\theta \left[\int_0^{2\pi} d\varphi e^{ir\Delta \sin\theta \cos\varphi} \cos\varphi \right]. \quad (26)$$

Writing $x = ir\Delta \sin\theta$ and changing $\varphi \mapsto \varphi + \pi$ the integral in the brackets in Eq. (26) becomes $2\pi i J_1(r\Delta \sin\theta)$. When this is inserted in Eq. (26) and steps analogous to the ones used in the evaluation of $\mathbf{J}_2(\Delta)$ are used, we get $\mathbf{K}_1(\Delta) = 0$.

III. CONCLUSIONS

We have shown through explicitly and detailed calculation that the integral of $\mathcal{G}(\mathbf{p})$ appearing in Eq. (42) of Ref. [4] is null for $\mathbf{x} - \mathbf{x}'$ lying in three orthonormal spatial directions in the rest frame of an arbitrary inertial frame $\mathbf{e}_0 = \partial/\partial t$.

This shows that no breakdown of locality concerning the anticommutator of $\{\Lambda(\mathbf{x}, t), \Pi(\mathbf{x}', t)\}$ and no preferred spacelike direction field in Minkowski spacetime is implied by the existence of Elko spinor fields.

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